

Bianchi cosmological models as integrable geodesic flows

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- Homogeneous cosmological models
- Geodesic flow on symmetric spaces

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Forty two years after the famous paper of

Francesco Calogero

*Solution of the one-dimensional n-body problem with quadratic
and/or inversely quadratic pair potentials*

Journal of Mathematical Physics 12 (1971) 419-436

Bill Sutherland

Physical Review A 1971, 1972

Jürgen Moser

Adv. Math. 1975

The **Calogero-Sutherland-Moser** type system, which consists of n -particles on a line interacting with pairwise potential $V(x)$ admits the following spin generalization

$$H_{ECM} = \frac{1}{2} \sum_{a=1}^n p_a^2 + \frac{1}{2} \sum_{a \neq b}^n f_{ab} f_{ba} V(x_a - x_b)$$

where

- I. $V(z) = z^{-2}$ Calogero
- II. $V(z) = a^2 \sinh^{-2} az$
- III. $V(z) = a^2 \sin^{-2} az$ Sutherland
- IV. $V(z) = a^2 \wp(az)$
- V. $V(z) = z^{-2} + \omega z^2$ Calogero

The nonvanishing Poisson bracket relations

$$\{x_i, p_j\} = \delta_{ij}, \quad \{f_{ab}, f_{cd}\} = \delta_{bc}f_{ad} - \delta_{ad}f_{cb}$$

The **Calogero-Sutherland-Moser** type system, which consists of n -particles on a line interacting with pairwise potential $V(x)$ admits the following spin generalization

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The configuration space \mathcal{S} of the real symmetric 3×3 matrices can be endowed with the flat Riemannian metric

$$ds^2 = \langle dQ, dQ \rangle = \text{Tr } dQ^2,$$

whose group of isometry is formed by orthogonal transformations $Q' = R^T Q R$. The system is invariant under the orthogonal transformations $S' = R^T S R$. The orbit space is given as a quotient space $\mathcal{S}/SO(3, \mathbb{R})$ which is a stratified manifold

(1) *Principal orbit-type stratum,*

when all eigenvalues are unequal $x_1 < x_2 < x_3$
with the smallest isotropy group $\mathbb{Z}_2 \otimes \mathbb{Z}_2$

(2) *Singular orbit-type strata*

forming the boundaries of orbit space with

(a) two coinciding eigenvalues (e.g. $x_1 = x_2$) ,
when the isotropy group is $SO(2) \otimes \mathbb{Z}_2$

(b) all three eigenvalues are equal ($x_1 = x_2 = x_3$) ,
here the isotropy group coincides with the isometry
group $SO(3, \mathbb{R})$

Let us consider the Hamiltonian system with the phase space spanned by the $N \times N$ symmetric matrices X and P with the noncanonical symplectic form

$$\{X_{ab}, P_{cd}\} = \frac{1}{2} (\delta_{ac}\delta_{bd} - \delta_{ad}\delta_{bc}) .$$

The Hamiltonian of the system defined as

$$H = \frac{1}{2} \mathbf{tr} P^2$$

describes a free motion in the matrix configuration space.

The following statement is fulfilled:

The Hamiltonian rewritten in special coordinates coincides with the Euler-Calogero-Moser Hamiltonian

$$H = \frac{1}{2} \sum_{i=1}^N p_i^2 + \frac{1}{2} \sum_{i \neq j}^N \frac{l_{ij}^2}{(x_i - x_j)^2}$$

with nonvanishing Poisson brackets for the canonical variables

$$\{x_i, p_j\} = \delta_{ij},$$

$$\{l_{ab}, l_{cd}\} = \frac{1}{2} (\delta_{ac}l_{bd} - \delta_{ad}l_{bc} + \delta_{bd}l_{ac} - \delta_{bc}l_{ad}) .$$

Let us introduce new variables

$$X = O^{-1}(\theta)Q(x)O(\theta),$$

where the orthogonal matrix $O(\theta)$ is parameterized by the $\frac{N(N-1)}{2}$ variables, e.g., the Euler angles $(\theta_1, \dots, \theta_{\frac{N(N-1)}{2}})$ and $Q = \text{diag}\|x_1, \dots, x_N\|$ denotes a diagonal matrix. This point

transformation induces the canonical one which we can obtain using the generating function

$$F_4 = \left[P, x_1, \dots, x_N, \theta_1, \dots, \theta_{\frac{N(N-1)}{2}} \right] = \mathbf{tr}[X(x, \theta)P].$$

Using the representation

$$P = O^{-1} \left[\sum_{a=1}^N \bar{\alpha}_a \bar{P}_a + \sum_{i < j=1}^{\frac{N(N-1)}{2}} \alpha_{ij} P_{ij} \right] O,$$

where the matrices $(\bar{\alpha}_a, \alpha_{ij})$ form an orthogonal basis in the space of the symmetric $N \times N$ matrices under the scalar

product

$$\text{tr}(\bar{\alpha}_a \bar{\alpha}_b) = \delta_{ab}, \quad \text{tr}(\alpha_{ij} \alpha_{kl}) = 2\delta_{ik}\delta_{jl}, \quad \text{tr}(\alpha_a \alpha_{ij}) = 0,$$

one can find that $\bar{P}_a = p_a$ and components P_{ab} are represented via the $O(N)$ right invariant vectors fields l_{ab}

$$P_{ab} = \frac{1}{2} \frac{l_{ab}}{x_a - x_b}.$$

The integration of the Hamilton equations of motion

$$\dot{X} = P, \quad \dot{P} = 0$$

derived with the help of Hamiltonian gives the solution of the Euler-Calogero-Moser Hamiltonian system as follows: *for the x -coordinates we need to compute the eigenvalues of the matrix $X = X(0) + P(0)t$, while the orthogonal matrix O , which diagonalizes X , determines the time evolution of internal variables.*

Forty six years after the famous papers of
Morikazu Toda

Journal of Physical Society of Japan 20 (1967) 431

Journal of Physical Society of Japan 20 (1967) 2095

The Hamiltonian of the non-periodic Toda lattice is

$$H_{NPT} = \frac{1}{2} \sum_{a=1}^n p_a^2 + g^2 \sum_{a=1}^{n-1} \exp [2(x_a - x_{a+1})],$$

and the Poisson bracket relations are

$$\{x_i, p_j\} = \delta_{ij}.$$

The equations of motion for the non-periodic Toda lattice are

$$\dot{x}_a = p_a, \quad a = 1, 2, \dots, n,$$

$$\dot{p}_a = -2 \exp [2(x_a - x_{a+1})] + 2 \exp [2(x_{a-1} - x_a)],$$

$$\dot{p}_1 = -2 \exp [2(x_1 - x_2)],$$

$$\dot{p}_n = -2 \exp [2(x_{n-1} - x_n)],$$

and for the periodic Toda lattice

$$\dot{x}_a = p_a,$$

$$\dot{p}_a = 2 \exp [2(x_{a-1} - x_a)] - \exp [2(x_a - x_{a+1})].$$

Further for simplicity we put $k_{ab} = 0$ and $\alpha = 0$.

We use the Gauss decomposition for the positive-definite symmetric matrix S

$$S = Z D Z^T.$$

Here $D = \text{diag} \|x_1, x_2, \dots, x_n\|$ is a diagonal matrix with positive elements and Z is an upper triangular matrix

$$Z = \begin{pmatrix} 1 & z_{12} & \dots & z_{1n} \\ 0 & 1 & \dots & z_{2n} \\ \vdots & \vdots & \ddots & \vdots \\ 0 & 0 & \dots & 1 \end{pmatrix}.$$

The differential of the symmetric matrix S is given by

$$dS = Z \left[dD + D\Omega + (D\Omega)^T \right] Z^T ,$$

where Ω is a matrix-valued right-invariant 1-form defined by

$$\Omega := dZZ^{-1} .$$

In the Lie algebra $gl(n, \mathbb{R})$ of $n \times n$ real matrices we introduce Weyl basis with elements

$$(e_{ab})_{ij} = \delta_{ai} \delta_{bj} ,$$

which are $n \times n$ matrices in the form

$$e_{ab} = \begin{pmatrix} & & & & & & \vdots & \\ & \dots & \dots & \dots & \dots & \dots & 1 & \dots \\ & & & & & & \vdots & \\ & & & & & & \vdots & \\ & & & & & & \vdots & \\ & & & & & & \vdots & \end{pmatrix}.$$

The scalar product in $gl(n, \mathbb{R})$ is defined by

$$(e_{ab}, e_{cd}) = \text{tr}(e_{ab}^T, e_{cd}) = \delta_{ac} \delta_{bd}.$$

Let us introduce also the matrices

$$\bar{\alpha}_a = \begin{pmatrix} 0 & \dots & \dots & \dots & 0 \\ \vdots & \ddots & & & \vdots \\ \vdots & & 1 & & \vdots \\ \vdots & & & \ddots & \vdots \\ 0 & \dots & \dots & \dots & 0 \end{pmatrix}$$

and

$$\alpha_{ab} = \begin{pmatrix} & & & \\ & \ddots & & \vdots \\ & & 1 & \dots \\ \dots & \dots & & \dots \\ & \vdots & & \vdots \\ & & 1 & \dots \\ & & \vdots & \vdots \end{pmatrix}.$$

with scalar product is defined by

$$(\bar{\alpha}_a, \bar{\alpha}_b) = \mathbf{tr}(\bar{\alpha}_a \bar{\alpha}_b) = \delta_{ab},$$

$$(\alpha_{ab}, \alpha_{cd}) = \mathbf{tr}(\alpha_{ab} \alpha_{cd}) = 2 (\delta_{ad} \delta_{bc} + \delta_{ac} \delta_{bd}),$$

$$(\bar{\alpha}_a, \alpha_{bc}) = \mathbf{tr}(\bar{\alpha}_a \alpha_{bc}) = 0.$$

Now we can write the differential of the matrix S in the form

$$dS = Z \left[\sum_{a=1}^n dx_a \bar{\alpha}_a + \sum_{a < b=1}^n x_a \Omega_{ab} \alpha_{ab} \right] Z^T.$$

Here Ω_{ab} are the coefficients of the matrix Ω

$$\Omega = \sum_{a < b} \Omega_{ab} e_{ab}.$$

In the case of Gauss decomposition of the symmetric matrix $S = ZDZ^T$ the corresponding momenta we seek in the form

$$P = (Z^T)^{-1} \left(\sum_{a=1}^n \bar{\mathcal{P}}_a \bar{\alpha}_a + \sum_{a < b} \mathcal{P}_{ab} \alpha_{ab} \right) Z^{-1}.$$

From the condition of invariance of the symplectic 1-form

$$\text{tr} (P dS) = \sum_{i=1}^n p_a dx_a + \sum_{a < b = 1}^n p_{ab} dz_{ab},$$

where the new canonical variables $(x_a, p_a), (z_{ab}, p_{ab})$ obey the Poisson bracket relations

$$\{x_a, p_b\} = \delta_{ab}, \quad \{z_{ab}, p_{cd}\} = \delta_{ac} \delta_{bd},$$

we obtain

$$\bar{\mathcal{P}}_a = p_a, \quad \mathcal{P}_{ab} = \frac{l_{ab}}{2x_a}.$$

Here l_{ab} are right-invariant vector fields on the group of upper triangular matrices with unities on the diagonal

$$l_{ab} = \left(\Omega^{-1} \right)_{ab,cd}^T p_{cd},$$

where $\Omega_{ab,cd}$ are coefficients in the decompositions of the 1-forms Ω_{ab} in coordinate basis

$$\Omega_{ab} = \Omega_{ab,cd} dz_{cd},$$

The canonical Hamiltonian in the new variables takes the form

$$H = \frac{1}{2} \sum_{a=1}^n p_a^2 x_a^2 + \frac{1}{4} \sum_{a < b=1}^n l_{ab}^2 \frac{x_a}{x_b}.$$

After the canonical transformation

$$x_a = e^{y_a}, \quad p_a = \pi_a e^{-y_a}$$

we arrive to Hamiltonian in the form of
spin nonperiodic Toda chain

$$H = \frac{1}{2} \sum_{a=1}^n \pi_a^2 + \frac{1}{4} \sum_{a < b=1}^n l_{ab}^2 e^{y_a - y_b}.$$

The generators of the group of upper triangular matrices are

$$l_{12} = p_{12} + \sum_{k=2}^n z_{2k} p_{1k}, \quad l_{13} = p_{13} + \sum_{k=2}^n z_{2k} p_{1k}.$$

The internal variables satisfy the Poisson bracket relations

$$\{l_{ab}, l_{cd}\} = C^{ef}_{ab,cd} l_{ef}$$

with structure coefficients

$$C^{ef}_{ab,cd} = \delta_{bc} \delta_{ae} \delta_{df}.$$

The generalized Toda chains are integrable systems associated with Lie algebras.

$$H = \frac{1}{2} \sum_{a=1}^n \pi_a^2 + \sum_{k=1}^r l_k^2 e^{2q_{\alpha_k}},$$

where $q_{\alpha_k} = \langle \alpha_k, q \rangle$.

Type A_{n-1}

The simple roots

$$\alpha_1 = e_1 - e_2, \dots, \alpha_{n-1} = e_{n-1} - e_n.$$

The potential

$$U = e^{q_1 - q_2} + \dots + e^{q_{n-1} - q_n}.$$

Type B_n

The simple roots

$$\alpha_1 = e_1 - e_2, \dots, \alpha_{n-1} = e_{n-1} - e_n, \alpha_n = e_n.$$

The potential

$$U = e^{q_1 - q_2} + \dots + e^{q_{n-1} - q_n} + e^{q_n}.$$

Type C_n

The simple roots

$$\alpha_1 = e_1 - e_2, \dots, \alpha_{n-1} = e_{n-1} - e_n, \alpha_n = 2e_n.$$

The potential

$$U = e^{q_1 - q_2} + \dots + e^{q_{n-1} - q_n} + e^{2q_n}.$$

Type D_n

The simple roots

$$\alpha_1 = e_1 - e_2, \dots, \alpha_{n-1} = e_{n-1} - e_n, \alpha_n = e_{n-1} + e_n.$$

The potential

$$U = e^{q_1 - q_2} + \dots + e^{q_{n-1} - q_n} + e^{q_{n-1} + q_n}.$$

Type G_2

The simple roots

$$\alpha_1 = e_1 - e_2, \alpha_2 = -2e_1 + e_2 + e_3.$$

The potential

$$U = e^{q_1 - q_2} + e^{-2q_1 + q_2 + q_3}.$$

Type F_4

The simple roots

$$\alpha_1 = e_1 - e_2, \alpha_2 = e_2 - e_3, \alpha_3 = e_3, \alpha_4 = 1/2(e_4 - e_1 - e_2 - e_3).$$

The potential

$$U = e^{q_1 - q_2} + e^{\frac{1}{2}(q_4 - q_1 - q_2 - q_3)}.$$

Type E_6

The simple roots

$$\begin{aligned}\alpha_1 &= 1/2(-e_1 + e_2 + \cdots + e_7 - e_8), \alpha_2 = e_1 - e_2, \\ \alpha_3 &= e_2 - e_3, \alpha_4 = e_3 - e_4, \alpha_5 = e_4 - e_5, \alpha_6 = -(e_1 + e_2).\end{aligned}$$

The potential

$$U = e^{q_1-q_2} + e^{q_2-q_3} + e^{q_3-q_4} + e^{q_4-q_5} + e^{-(q_1+q_2)} + \\ + e^{\frac{1}{2}(-q_1+q_2+q_3+q_4+q_5+q_6+q_7-q_8)}.$$

Homogeneous cosmological models

Spacetime decomposition

Let (\mathcal{M}, g) be a smooth four-dimensional paracompact and Hausdorff manifold. In each point of the open set $\mathcal{U} \subset \mathcal{M}$ we propose that are defined the local basis of 1-forms and its dual basis

$$\{\omega^\mu, \mu = 0, 1, 2, 3\} \quad \{e_\nu, \nu = 0, 1, 2, 3\}$$

such that

$$[e_\alpha, e_\beta] = C^\mu{}_{\alpha\beta} e_\mu,$$

where $C^\mu{}_{\alpha\beta}$ are the basis structure functions. The symmet-

ric metric is

$$g = g_{\mu\nu} \omega^\mu \otimes \omega^\nu, \quad g^{-1} = g^{\mu\nu} e_\mu \otimes e_\nu$$

We suppose that on the manifold (\mathcal{M}, g) is defined affine connection ∇

$$\Gamma^\mu{}_\nu = \Gamma^\mu{}_{\nu\alpha} \omega^\alpha \iff \Gamma^\mu{}_{\beta\alpha} = \langle \omega^\mu, \nabla_{e_\alpha} e_\beta \rangle$$

In manifold with affine connection we can construct the

bilinear antisymmetric mapping

$$T(X, Y) = \nabla_X Y - \nabla_Y X - [X, Y]$$

The First Cartan structure equation is

$$d\omega^\mu + \Gamma^\mu{}_\alpha \wedge \omega^\alpha = \frac{1}{2} T^\mu$$

The tensor type (1, 3) defined by

$$R(\omega, Z, X, Y) = \langle \omega, R(X, Y)Z \rangle$$

is called the curvature tensor, where

$$R(X, Y) = \nabla_X \nabla_Y - \nabla_Y \nabla_X - [\nabla_X, \nabla_Y]$$

If we define the curvature 2-form $\Omega^\mu{}_\nu$ by

$$\Omega^\mu{}_\nu = d\Gamma^\mu{}_\nu + \Gamma^\mu{}_\alpha \wedge \Gamma^\alpha{}_\nu$$

one can obtain the Second structure Cartan equation

$$\Omega^\mu{}_\nu = \frac{1}{2} R^\mu{}_{\nu\alpha\beta} \omega^\alpha \wedge \omega^\beta$$

In the Riemannian geometry the commutator of two vector fields is given by

$$[X, Y] = \nabla_X Y - \nabla_Y X$$

and for every vector field X we have the condition $\nabla_X g = 0$.

The connection and the Riemann tensor in the noncoordinate basis $\{e_\nu\}$ are given with

$$\Gamma^\mu_{\alpha\beta} = \frac{1}{2}g^{\mu\sigma}(e_\alpha g_{\beta\sigma} + e_\beta g_{\alpha\sigma} - e_\sigma g_{\alpha\beta})$$

$$-\frac{1}{2}g^{\mu\sigma}(g_{\alpha\rho}C^\rho_{\beta\sigma} + g_{\beta\rho}C^\rho_{\alpha\sigma}) - \frac{1}{2}C^\mu_{\alpha\beta},$$

$$R^\sigma_{\alpha\mu\nu} = e_\mu\Gamma^\sigma_{\alpha\nu} - e_\nu\Gamma^\sigma_{\alpha\mu} + \Gamma^\rho_{\alpha\nu}\Gamma^\sigma_{\rho\mu} - \Gamma^\rho_{\alpha\mu}\Gamma^\sigma_{\rho\nu} - \Gamma^\sigma_{\alpha\rho}C^\rho_{\mu\nu},$$

where $C^\mu_{\alpha\beta} = \Gamma^\mu_{\beta\alpha} - \Gamma^\mu_{\alpha\beta}$.

Hereinafter we suppose that

$$\mathcal{M} = T_1 \times \Sigma_t$$

Let us introduce the noncoordinate basis of vector fields $\{e_\perp, e_a\}$

$$[e_\perp, e_a] = C^\perp{}_{\perp a} e_\perp + C^d{}_{\perp a} e_d,$$

$$[e_a, e_b] = C^d{}_{ab} e_d,$$

with the structure functions

$$C^\perp{}_{\perp a} = e_a \ln N, \quad C^d{}_{\perp a} = \frac{1}{N} (N^b C^d{}_{ab} + e_a N^d)$$

The metric in the corresponding dual basis $\{\theta^\perp, \theta^a\}$

$$g = -\theta^\perp \otimes \theta^\perp + \gamma_{ab} \theta^a \otimes \theta^b,$$

where γ is the induced metric on the submanifold Σ_t . The components of the vector field X_0 in this basis

$$X_0 = Ne_{\perp} + N^a e_a$$

are the Lagrange multipliers in ADM scheme which can be obtained if we pass to the coordinate basis $\{X_0 = \frac{\partial}{\partial t}, \frac{\partial}{\partial x^a}\}$

$$g = - \left(N^2 - N^a N_a \right) dt \otimes dt + 2N_a dt \otimes dx^a + \gamma_{ab} dx^a \otimes dx^b$$

The second fundamental form characterizes the embedding of

Σ_t in (\mathcal{M}, g)

$$K(X, Y) : T_x \mathcal{M} \times T_x \mathcal{M} \rightarrow \mathbb{R}$$

$$K(X, Y) = \frac{1}{2}(Y \cdot \nabla_X e_\perp - X \cdot \nabla_Y e_\perp), \quad X, Y \in T\Sigma_t$$

The another representation for the extrinsic curvature is

$$K_{ab} = -\frac{1}{2}\mathcal{L}_{e_\perp} \gamma_{ab}$$

To find the $3 + 1$ decomposition we define the induced on

the Σ_t connection

$${}^3\Gamma^c_{ba} := \Gamma^c_{ba} = \langle {}^3\theta^c, {}^3\nabla_{e_a} e_b \rangle,$$

where ${}^3\nabla_X$ is the covariant derivative with respect to γ and corresponding curvature tensor is

$${}^3R(X, Y) = {}^3\nabla_X {}^3\nabla_Y - {}^3\nabla_Y {}^3\nabla_X - [{}^3\nabla_X, {}^3\nabla_Y]$$

In the basis $\{e_\perp, e_a\}$ we find the components of the connection

$$\begin{aligned}
\Gamma^\perp_{\perp\perp} &= 0, & \Gamma^\perp_{\perp i} &= 0, \\
\Gamma^j_{\perp\perp} &= c^j, & \Gamma^j_{\perp i} &= -K^j_i, \\
\Gamma^\perp_{i\perp} &= c_i, & \Gamma^\perp_{ji} &= -K_{ji}, \\
\Gamma^j_{i\perp} &= -K^j_i + \langle {}^3\theta^j, \mathcal{L}_{e_\perp} e_i \rangle, & \Gamma^k_{ij} &= {}^3\Gamma^k_{ij}
\end{aligned}$$

and of the Riemann tensor

$$R^j_{\perp k\perp} = K^{js} K_{sk} + \gamma^{js} \mathcal{L}_{e_\perp} K_{ks} + \frac{1}{N} \gamma^{js} {}^3\nabla_{e_k} {}^3\nabla_{e_s} N,$$

$$R^\perp_{ijk} = {}^3\nabla_{e_k} K_{ij} - {}^3\nabla_{e_j} K_{ik},$$

$$R^s_{ijk} = {}^3R^s_{ijk} + K_{ik} K_j{}^s - K_{ij} K_k{}^s$$

Finally the scalar curvature can be obtained in the form

$$R = {}^3R + K_a{}^a K_b{}^b - 3K_{ab}K^{ab} - \frac{2}{N}\gamma^{ab}{}^3\nabla_{e_a}{}^3\nabla_{e_b}N - 2\gamma^{ab}\mathcal{L}_{e_\perp}K_{ab}$$

The Hilbert-Einstein action is

$$\begin{aligned} L[N, N_a, \gamma_{ab}, \dot{N}, \dot{N}_a, \dot{\gamma}_{ab}] &= \iint_{t_1S}^{t_2} dt \, {}^3\sigma N \left\{ {}^3R + K_a{}^a K_b{}^b - 3K_{ab}K^{ab} \right\} \\ &\quad - \iint_{t_1S}^{t_2} dt \, {}^3\sigma N \left\{ \frac{2}{N}\gamma^{ab}{}^3\nabla_{e_a}{}^3\nabla_{e_b}N - 2\gamma^{ab}\mathcal{L}_{e_\perp}K_{ab} \right\} \end{aligned}$$

where ${}^3\sigma = \sqrt{\gamma} \theta^1 \wedge \theta^2 \wedge \theta^3$ is the volume 3-form on Σ_t .

Bianchi models

L. Bianchi 1897

Sugli spazii atre dimensioni du ammettono un gruppo continuo di movimenti

Soc. Ital. della Sci. Mem. di Mat.
(Dei. XL) (3) 3 267

By definition, Bianchi models are manifolds with product topology

$$\mathcal{M} = \mathbb{R} \times G_3$$

On the three dimensional Riemannnian manifold $\Sigma_t\gamma$ there

exist left-invariant 1-forms $\{\chi^a\}$ such that

$$d\chi^a = -\frac{1}{2}C^a_{bc}\chi^b \wedge \chi^c$$

The dual vector fields $\{\xi_a\}$ form a basis in the Lie algebra of the group G_3

$$[\xi_a, \xi_b] = C^d_{ab}\xi_d$$

with structure constants $C^d_{ab} = 2d\chi^d(e_a, e_b)$. In the invariant basis

$$[e_\perp, e_a] = C^d_{\perp a}e_d, \quad [e_a, e_b] = -C^d_{ab}e_d,$$

with $C^d_{\perp a} = N^{-1}N^b C^d_{ab}$ the metric takes the form

$$g = -\theta^\perp \otimes \theta^\perp + \gamma_{ab}\theta^a \otimes \theta^b$$

The preferable role of this choice for a coframe is that the functions N, N^a and γ_{ab} depend on the time parameter t only. Due to this simplification the initial variational problem for Bianchi A models is restricted to a variational problem of the “mechanical” system

$$L(N, N_a, \gamma_{ab}, \dot{\gamma}_{ab}) = \int_{t_1}^{t_2} dt \sqrt{\gamma} N \left[{}^3R - K_a{}^a K_b{}^b + K_{ab} K^{ab} \right],$$

where 3R is the curvature scalar formed from the spatial metric γ

$${}^3R = -\frac{1}{2}\gamma^{ab}C_{da}^c C_{cb}^d - \frac{1}{4}\gamma^{ab}\gamma^{cd}\gamma_{ij}C^i{}_{ac}C^j{}_{bd},$$

Table 1

The Bianchi-Behr classification of groups

class	type	N			a
A	I	0	0	0	0
	II	1	0	0	0
	VI ₀	0	1	-1	0
	VII ₀	0	1	1	0
	VIII	1	1	-1	0
	IX	1	1	1	0

Table 2

The Bianchi-Behr classification of groups

	V	0	0	0	1
	IV	0	0	1	1
B	VI _a , $a < 0$ (III=VI ₋₁)	0	1	-1	$\sqrt{-a}$
	VII _a , $a > 0$	0	1	1	\sqrt{a}

and

$$K_{ab} = -\frac{1}{2N} \left((\gamma_{ad} C^d_{bc} + \gamma_{bd} C^d_{ac}) N^c + \dot{\gamma}_{ab} \right)$$

Table 3

Relation between Thurston's geometries and BKS types

Thurston's geometries	BKS types	class	sectional curvature
\mathbb{R}^3	I, VII_0	A	0, 0, 0
\mathbb{S}^3	IX	A	1, 1, 1
\mathbb{H}^3	V	B	-1, -1, -1
	$\text{VII}_a, a > 0$		$-a^2, -a^2, -a^2$
$\mathbb{S}^2 \times \mathbb{R}$	◀◀ ▲ 55/67 ▶▶	1, 2, 2	

Table 4

Relation between Thurston's geometries and BKS types

$\widetilde{SL(2)\mathbb{R}}$	VIII	A	$-\frac{5}{4}, -\frac{1}{4}, -\frac{1}{4}$
Nil	II	A	$-\frac{1}{4}, \frac{1}{4}, \frac{1}{4}$
Sol	VI_0	A	1, -1, -1

is the extrinsic curvature of the slice Σ_t defined by the

relation

$$K_{ab} = -\frac{1}{2}\mathcal{L}_{e_\perp}\gamma_{ab}$$

In the theory we have four primary constraints

$$\pi := \frac{\delta L}{\delta \dot{N}} = 0, \quad \pi^a := \frac{\delta L}{\delta \dot{N}_a} = 0$$

$$\pi^{ab} := \frac{\delta L}{\delta \dot{\gamma}_{ab}} = \sqrt{\gamma}(\gamma_{ab}K^i{}_i - K^{ab})$$

The symplectic structure on the phase space is defined by

the following non-vanishing Poisson brackets

$$\{N, \pi\} = 1, \quad \{N_a, \pi^b\} = \delta_a^b, \quad \{\gamma_{cd}, \pi^{rs}\} = \frac{1}{2} \left(\delta_c^r \delta_d^s + \delta_c^s \delta_d^r \right)$$

Due to the reparameterization symmetry of inherited from the diffeomorphism invariance of the initial Hilbert-Einstein action, the evolution of the system is unambiguous and it is governed by the total Hamiltonian

$$H_T = N\mathcal{H} + N^a\mathcal{H}_a + u_0P^0 + u_aP^a$$

with four arbitrary functions $u_a(t)$ and $u_0(t)$. One can verify

that the secondary constraints are first class and obey the algebra

$$\{\mathcal{H}, \mathcal{H}_b\} = 0, \quad \{\mathcal{H}_a, \mathcal{H}_b\} = -C^d{}_{ab}\mathcal{H}_d$$

From the condition of time conservation of the primary constraints follows four secondary constraints

$$\mathcal{H} = \frac{1}{\sqrt{\gamma}} \left(\pi^{ab} \pi_{ab} - \frac{1}{2} \pi^a{}_a \pi^b{}_b \right) - \sqrt{\gamma} {}^3R, \quad \mathcal{H}_a = 2 C^d{}_{ab} \pi^{bc} \gamma_{cd},$$

which obey the algebra

$$\{\mathcal{H}, \mathcal{H}_a\} = 0, \quad \{\mathcal{H}_a, \mathcal{H}_b\} = -C^d{}_{ab}\mathcal{H}_d$$

The Hamiltonian form of the action for the Bianchi A models can be obtained in the form

$$L[N, N_a, \gamma_{ab}, \pi, \pi^a, \pi^{ab}] = \int_{t_1}^{t_2} \pi^{ab} d\gamma_{ab} - H_C dt$$

where the canonical Hamiltonian is a linear combination of the secondary constraints $H_C = N\mathcal{H} + N^a\mathcal{H}_a$.

Hamiltonian reduction of Bianchi I cosmology

In the case of Bianchi I model the group which acts on (Σ_t, γ) is T^3 and the action takes the form

$$L[N, N_a, \gamma_{ab}, \pi, \pi^a, \pi^{ab}] = \int_{t_1}^{t_2} \pi^{ab} d\gamma_{ab} - N \mathcal{H} dt.$$

Using the decomposition for arbitrary symmetric non-singular matrix

$$\gamma = R^T(\chi) e^{2X} R(\chi),$$

where $X = \text{diag}\|x_1, x_2, x_3\|$ is diagonal matrix and

$$R(\psi, \theta, \phi) = e^{\psi J_3} e^{\theta J_1} e^{\phi J_3}$$

we can pass to the new canonical variables

$$(\gamma_{ab}, \pi^{ab}) \implies (\chi_a, p_{chi_a}; x_a, p_a)$$

The corresponding momenta are

$$\pi = R^T \left(\sum_{s=1}^3 \bar{\mathcal{P}}_s \bar{\alpha}_s + \sum_{s=1}^3 \mathcal{P}_s \alpha_s \right) R$$

where

$$\bar{\mathcal{P}}_a = p_a,$$

$$\mathcal{P}_a = -\frac{1}{4} \frac{\xi_a}{\sinh(x_b - x_c)}, \quad (\text{cyclic permutations } a \neq b \neq c)$$

and the left-invariant basis of the action of the $SO(3, \mathbb{R})$ in the phase space with three dimensional orbits is given by

$$\xi_1 = \frac{\sin \psi}{\sin \theta} p_\phi + \cos \psi p_\theta - \sin \psi \cot \theta p_\psi,$$

$$\xi_2 = -\frac{\cos \psi}{\sin \theta} p_\phi + \sin \psi p_\theta + \cos \psi \cot \theta p_\psi ,$$

$$\xi_3 = p_\psi$$

In the new variables the Hamiltonian constraint reads

$$\mathcal{H}_{BI} = \frac{1}{2} \sum_{a=1}^3 p_a^2 - \sum_{a < b} p_a p_b + \frac{1}{2} \sum_{(abc)} \frac{\xi_c^2}{\sinh^2(x_a - x_b)}$$

Another form for the Bianchi I Hamiltonian is

$$\mathcal{H}_{BI} = \frac{1}{2} \sum_{a=1}^3 p_a^2 - \sum_{a < b} p_a p_b + \frac{1}{4} \sum_{a < b=1}^3 l_{ab}^2 e^{y_a - y_b} .$$

Hamiltonian reduction of Bianchi II cosmology

By definition on the three dimensional spacelike submanifold acts the Heisenberg group $H(1)$. The right-invariant vector fields we choose to satisfy the Poisson bracket relations

$$\{\xi_1^R, \xi_2^R\} = \xi_3^R,$$

$$\{\xi_1^R, \xi_3^R\} = 0,$$

$$\{\xi_2^R, \xi_3^R\} = 0.$$

If we use the Gauss decomposition for the 3-metric

$$\gamma = Z^T D Z$$

we obtain the Hamiltonian constraint for the Bianchi II model in the form

$$\mathcal{H}_{BII} = p_1 + p_2 + p_3 - \frac{1}{2} (p_1 + p_2 + p_3) + \frac{1}{2} l_{23}^2 e^{y_3 - y_2} + \frac{1}{2} e^{2y_1}.$$

Many thanks

TO THE ORGANIZERS

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